

High spatial Resolution mapping of Venus Mesospheric Winds by infrared heterodyne Spectroscopy of CO₂

G. Sonnabend (1), **M. Sornig** (1), P. Krötz (1), D. Stupar (1), T. Livengood (2), R. Schieder (1), T. Kostiuik (2)

(1) I. Physikalisches Institut, University of Cologne, (2) NASA Goddard Space Flight Center, Greenbelt MD, USA, (samstag@ph1.uni-koeln.de)

We present wind measurements in the Venusian upper mesosphere / lower thermosphere by means of infrared heterodyne spectroscopy of CO₂ features at 959.3917 cm⁻¹. Observations are carried out using the Cologne Tuneable Heterodyne Infrared Spectrometer (THIS) from May 25th to June 6th 2007 shortly before Venus superior conjunction at the McMath-Pierce solar telescope on Kitt Peak in Arizona.

Providing high spectral resolution winds can be retrieved from Doppler-shifts of CO₂ non-thermal emission from the upper mesosphere. The sub-solar to anti-solar flow (SS-AS flow) and the retrograde superrotating zonal circulation (RSZ) are targeted and observations are carried out systematically on the day-side of the planet which is illuminated appr. 50%. The mesospheric region is of special interest because it is the not very well understood transitions zone from the RSZ dominated troposphere and the SS-AS flow dominated thermosphere.

Measurements are part of the coordinated ground-based observing campaign to support VenusExpress from May 25th to June 9th. Complementary ground based observing methods probing wind velocities at different altitudes in the atmosphere of Venus provide the possibility to get a vertical wind profile. E.g. Doppler shifts of CO₂ lines at visible wavelength together with reflected solar Fraunhofer lines probe dynamics at the cloud tops and a few kilometer above while interferometric CO millimeter observations provide information about the lower mesosphere and sub-millimeter spectral line observations are pointing to a region between 95 and 105km. The presented mid-IR measurements probe an altitude of 100-120km.