



The SECCHI experiment on the STEREO mission

R. Howard (1), D. Moses (1), A. Vourlidas (1), J. Newmark (1) D. Socker (1), D. Wang (1), R. Baugh (1), D. McMullin (1), J. Davila (2), W. Thompson (2), B. Klein (2), J. Lemen (3), J-P Wuelsel (3), R. Harrison (4), N. Waltham (4), J. Lang (4), C. Eyles (5), J-M Defise (6), J-P Halain (6), V. Bothmer (7), J-P Delaboudiniere (8), F. Auchere (8), R. Mercier (9), M-F Ravet (9)

(1) Code 7660, U.S. Naval Research Lab, Washington DC 20375, USA (2) Code 682, NASA/GSFC, Greenbelt MD 20771, USA (3) Lockheed Martin Solar & Astrophysics Lab, 3251 Hanover Street, Palo Alto, CA 94304, USA (4) CCLRC Rutherford Appleton Laboratory, Chilton Didcot, OX11 0QX, UK, (5) School of Physics and Astronomy, University of Birmingham, Edgbaston, Birmingham, UK B15 2TT (6) Centre Spatial de Liege, Avenue du Pre Aily, B-4031 Angleur, Belgium, (7) Institut für Astrophysik, Friedrich-Hund-Platz, Universität Göttingen, 37077 Göttingen, Germany, (8) Institute d' Astrophysique Spatiale, Université Paris - Sud / CNRS, F-91405 Orsay, France, (9) Laboratoire Charles Fabry, Institute d'Optique, F-91403 Orsay, France

The Sun Earth Connection Coronal and Heliospheric Investigation (SECCHI) on the NASA Solar Terrestrial Relations Observatory (STEREO) mission is a suite of remote sensing instruments consisting of an extreme ultraviolet (EUV) imager, two white light coronagraphs, and two telescopes that comprise the heliospheric imager. SECCHI will observe coronal mass ejections (CMEs) from their birth at the sun, through the corona and into the heliosphere. A complete instrument suite is being carried on each of the two STEREO spacecraft, which will provide the first sampling of a CME from two vantage points. The spacecraft, launched 25 October 2006, are orbiting the Sun, one Ahead of the Earth and the other Behind, each separating from Earth at about 22 degrees per year. The varying separation means that we will have different observational capabilities as the spacecraft separate and therefore differing science goals. The primary science objectives all are focused on understanding the physics of the CME process - their initiation, 3D morphology, propagation, interaction with the interplanetary medium and space weather effects. By observing the CME from multiple viewpoints with UV and coronagraphic telescopes and by combining these observations with ra-

dio and in-situ observations from the other instruments on STEREO as well as from other satellites and ground based observatories operating at the same time, answers to some of the outstanding questions will be obtained. All of the telescopes are working very well and have been producing spectacular images. We will show examples of some of the data and some of the initial results.