



## **Expected Observations of the Active Solar X-ray Corona at 6-41Å with the ATSSI Microcalorimeter Sounding Rocket Payload**

**D. S. Martínez-Galarce** (1), P. F. Boerner (1), S. Deiker (1), B. Cabrera (2), D. Chakraborty (2), R. Kumar (2), T. W. Barbee, II (3), K. Irwin (4) and P. C. Baker (5)

(1) Lockheed Martin Solar & Astrophysics Laboratory, California USA (denmart@lmsal.com / Fax: (650) 424-3994) / Phone: (650) 354-5008) (2) Stanford University, California USA (3) Lawrence Livermore National Laboratory, California USA (4) National Institute of Standards & Technology, Colorado USA (5) Baker Consulting, California USA

The Advanced Technology Solar Spectroscopic Imager (ATSSI) sounding rocket payload will employ a new technology permitting true spectroheliograms to be recorded in the  $\sim 6-41 \text{ \AA}$  ( $300\text{eV}-2 \text{ keV}$ ) range. Implementation of this novel class of 3-D detector (in Energy, Time and Position), a *quantum microcalorimeter*, placed at the focus of a Wolter I grazing-incidence telescope, is the principal goal of the mission. Over a nominal 300 second period of observation, and with a cadence of  $\leq 10 \text{ sec}$ , a  $2 \times 2$  array of detectors with  $\sim 5 \text{ arc sec}$  pixels, will record  $\sim 120$  spectra ( $\sim 5000 \text{ cts}$  each) with spectral resolution of  $\sim 3\text{eV}$  (at  $2\text{keV}$ , i.e.  $E/\Delta E \sim 667$ ). Each spectrum will include emission lines of ions ranging from C V to S XVI, generated over the temperature range from  $\sim 600,000 \text{ K}$  to  $\sim 6,300,000 \text{ K}$ . Using these observations, the goal of the investigation will be to measure and determine non-thermal episodic heating mechanism(s) in coronal loops.