Long-term variations in the properties of large solar energetic particle events.

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The occurrence of large solar energetic particle (SEP) events (30MeV fluence $>2x10^9$ cm^{-2}) is reviewed, using data obtained by satellites, ground based instrumentation, and from the glaciological record. It is shown that the frequency of occurrence of large fluence events during the space era, 1960 et seq, was one of the lowest in the past 400 years. Large SEP were up to 6 to 8-times more frequent during periods of low solar activity, such as the Gleissberg minimum, circa 1890-1900. The possibility that the frequency of occurrence of SEP is inversely related to the strength of the magnetic fields near the Sun is discussed, and advanced as a working hypothesis that allows the occurrence of SEP during the 11-year solar cycle, and over centennial scales, to be better understood. The energy dependence of the particle spectra at high energies; the occurrence of large ephemeral anisotropies; and the dependence of the particle intensities upon heliographic longitude are summarized. The "Carrington white light flare" of 1859 resulted in the greatest 30MeV fluence $(18.8 \times 10^9 \text{ cm}^{-2})$ near Earth in the past 400 years, however, evidence from the ¹⁰Be record might indicate a significantly greater event during the Spoerer "grand minimum" of the Sun (1420-1540). A cumulative probability curve for the occurrence of very large SEP events is presented. In closing, a prediction is made regarding the occurrence of large SEP events until 2050AD.