

# **The correlation between radio-luminosity and spindown energy loss in pulsars**

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It is conventionally suggested that the beamed radio emission of pulsars could have three components: core, inner-cone, and outer-cone. A pulse-profile observed depends on the geometry of an observer and the star (i.e., inclination angle and viewing angle), and the intrinsic radio-luminosities of pulsars should accordingly not be the detected ones that are averaged over an emission phase. Based on a large sample with high quality, a relationship between intrinsic radio-luminosity and rotation-energy loss rate is obtained, with the inclusion of the geometrical effects, for both normal and millisecond pulsars. An acceptable pulsar radio-emission model should certainly be able to reproduce this correlation.