Simulation of the Lunar Surface Emission and Inversion of the Regolith Layer Thickness Using Fusio

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A correspondence of the lunar regolith layer thickness to the lunar digital elevation mapping (DEM) is p to construct the global distribution of lunar regolith layer thickness. Based on some measurements, the temperature distribution over the lunar surface is proposed. Albedo of the lunar nearside at the way 0.42, 0.65, 0.75, 0.95  $\mu$ m from the telescopic observation is employed to construct the spatial distributi FeO+TiO<sub>2</sub>on the lunar regolith layer. A statistic relationship between the DEM and FeO+TiO<sub>2</sub> contec lunar nearside is then extended to construction of FeO+TiO<sub>2</sub> content of the lunar farside. Thus, the opermittivity of global lunar regolith layer can be determined.

Based on all theses conditions, brightness temperature of the lunar regolith layer in passive microwav sensing, which is planned for China's Chang-E lunar project, is numerically simulated by a parallel lay using the fluctuation dissipation theorem.

Furthermore, taking these simulations as observations, an inversion method of the lunar regolith layer t is developed by using three- or two-channels brightness temperatures. When the FeO+TiO<sub>2</sub> conter and the four channels brightness temperatures in Chang-E project are well distinguishable, the regol thickness and physical temperature of the underlying lunar rocky media can be inverted by the threeapproach. When the FeO+TiO<sub>2</sub> content is so high that the brightness temperature at higher frequency might be saturated, the regolith layer thickness is alternatively inverted only by the two-channels appro-Numerical simulation and inversion approach in this paper make an evaluation of the performance passive microwave remote sensing, and for future data calibration and validation.