Geophysical Research Abstracts, Vol. 7, 10270, 2005 SRef-ID: 1607-7962/gra/EGU05-A-10270 © European Geosciences Union 2005



Thermal Structure and Ion-Neutral Coupling at Jupiter

T. Majeed (1,2), J. H. Waite (1), S. W. Bougher (1) and G. R. Gladstone (3) (1) University of Michigan (tariqm@umich.edu), (2) American University of Sharjah, UAE (tmajeed@sharjah.edu), (3) Southwest Research Institute

Recent studies of high-latitude ionosphere-magnetosphere coupling at Jupiter have indicated that the collisions between ions, magnetically connected to sub-rotating regions of the magnetosphere, and neutral species produces a rotational slippage of the neutral atmosphere. Such coupling is introduced via an ion-neutral coupling parameter, K, which is defined as a measure of reductions in the Pedersen conductivity as the corotation breaks down in the Jovian magnetosphere. While no *in-situ* measurement is available for the neutral atmosphere of Jupiter's auroral region, infrared and ultraviolet spectrographic imaging results indicate auroral exospheric temperatures > 1600 K. Furthermore, high-resolution infrared and far ultraviolet spectroscopy of the aurora suggests the presence of high-speed (> 2 km/s) winds in the Jovian thermosphere.

We use our fully coupled 3-D Jupiter Thermosphere General Circulation Model (JT-GCM) from 20 μ bar to 10^{-4} nbar to address above issues on a global scale. Such general circulation models allow the global dynamical structure to be simulated self-consistently with the thermal structure and compositions (ion and neutral). The coupling between ions in the Jovian auroral ovals and the corotating neutral atmosphere can also be simulated. The heat sources that drive the thermospheric flow are due to solar EUV radiation and high-latitude auroral processes such as particle precipitation and Joule heating. Simulations of Jupiter's global thermospheric dynamics indicate significant ion transport by high-speed winds. Strong neutral outflows also develop with velocities >1 km/s and temperatures up to 4000 K (depending on the magnitude of Joule heating). The models demonstrate that a significant amount of auroral energy is transported to equatorial latitudes. Auroral temperatures inferred from remote sensing experiments can be explained. The values of K are obtained for standard model inputs to understand the characteristics of the ion-neutral coupling at Jupiter.