



Solar wind and IMF control of substorm onset

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Few studies on solar wind control of substorm onset exist and only two have good statistics using the IMF. Those studies with solar wind statistics employ only a few 10s of events, yet it has been proposed that solar wind dynamic pressure changes may trigger substorms. While a small percentage of substorms are believed to be triggered by solar wind dynamic pressure variations and hence, solar wind control of the substorms is not surprising, a significant number of substorm events exist that appear to have either IMF triggers or no triggers at all. We will show the results of a statistical examination of solar wind control of substorm onset with a database of over 3300 onsets using solar wind data propagated with the Weimer et al. [2003] technique. Our preliminary work demonstrates that the magnetic latitude of the substorm onset decreases with increasing magnitude of solar wind dynamic pressure, IMF B_x, B_y, and B_z. The solar wind plasma or IMF does not appear to influence the substorm onset magnetic local time. Furthermore, substorm onset location does not appear to be a function of dipole tilt, season, and universal time as one might expect. We will also present results examining the influence of triggered and non-triggered substorms as well as isolated and non-isolated substorms. The results of this study will be of interest to the upcoming THEMIS mission.