



## **Energetic ions upstream from the saturnian magnetosphere: Cassini/MIMI observations.**

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The Magnetospheric Imaging Instrument (MIMI) performed comprehensive measurements of the energetic ion population in the environment upstream from the dawn side of the Saturnian Magnetosphere during the approach phase and the subsequent first orbit of the Cassini spacecraft around the planet. High sensitivity observations of energetic ion directional intensities, energy spectra and ion composition were obtained by the Ion and Neutral Camera (INCA) of the MIMI Instrument complement with a geometry factor of  $\sim 2.5 \text{ cm}^2 \text{ sr}$  and some capability of separating light (H, He) and heavier (C, N, O) ion groups (henceforth referred to as “hydrogen” and “oxygen” respectively). The observations have revealed the presence of a series of distinct upstream bursts of energetic hydrogen and oxygen ions up to distances of 126 Rs, which exhibit the following characteristics: (1) The hydrogen ion bursts are observed in the energy range  $\sim 3$  to  $\sim 220$  keV (and occasionally to  $E > 220$  keV) and the oxygen ion bursts in the energy range  $\sim 32$  to  $\sim 300$  keV. (2) The duration of the ion bursts is several minutes up to 4 hrs. (3) The events are of varying composition, with some exhibiting significant fluxes of oxygen. (4) The bursts have a filamentary structure with some exhibiting distinct signatures of “velocity-filtering effects” at the edges of convecting IMF filaments. (5) Some ion bursts are accompanied by distinct diamagnetic field depressions. Furthermore there are indications that the ambient energetic ion fluxes appear to be modulated with the planetary period (11 hrs) at large distances

( $\sim 800 R_S$ ) from Saturn. Given that energetic ions trapped within the magnetosphere of Saturn are mostly  $H^+$  and  $O^+$  (Krimigis et al, 2005), we conclude that  $O^+$ -rich upstream events must be particles leaking from Saturn's magnetosphere under favorable IMF conditions. The observations will be presented and compared to theoretical models.

Reference:

Krimigis, S.M, et al, Science, 307 (in press), 2005