

Particle simulation of resonant interactions of electrons with chorus emissions in a nonuniform magnetic field

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The GEOTAIL spacecraft has been observing many chorus emissions in the dayside outer geomagnetosphere with the Plasma Wave Instrument (PWI). It is generally considered that the chorus emissions are generated via nonlinear wave-particle interaction between anisotropic electrons and whistler mode waves. With the well-known linear growth rate of a parallel-propagating whistler mode wave, pitch-angle anisotropy and its critical value are an important factor in determining whether a whistler mode waves will grow through cyclotron resonance with resonant electrons or not. In this study, we analyze the wave-particle interaction involved in the generation and propagation of the chorus emissions by using electromagnetic particle simulation. In the simulation, we inject the whistler mode wave parallel to the magnetic field into hot electrons with pitch-angle anisotropy in a one-dimensional model. From this result, only the wavefront (several wavelength) of the injected whistler mode wave grows, and the wave propagating behind the amplified portion does not grow. This suggests that the resonant electrons with large pitch-angle anisotropy are pitch-angle diffused by the cyclotron resonance between the whistler mode wavefront and the resonant electrons. As a result, the initially unstable pitch-angle anisotropy is lost, and such electrons can no longer grow the wave behind the wavefront, and the wavefront saturates. We also evaluate pitch-angle anisotropies and linear growth rates from the electron velocity distributions with the cyclotron resonance. By the cyclotron interaction, the initially unstable resonant electrons with the velocity corresponding to the whistler mode wave resonant frequency are diffused. Then the anisotropy decreases down to the stable anisotropy, and the linear growth rate is reduced to 0. Next, to simulate generation and propagation process of chorus emissions around the equatorial plane in the magnetosphere, we tentatively set up a spatially inhomogeneous magnetic field as the Earth's dipole magnetic field in the simulation, and simulate a process that whistler mode waves triggered from a thermal noise existing in space are generated and propagate. From the simulation result, waves are generated in the region where the intensity of the magnetic field is weak. This is consistent with the theory of chorus generation and chorus observation around the equatorial plane. We will present the evolution of the resonant electrons in the inhomogeneous magnetic field.